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THE EQUATION OF STATE OF BERYLLIUM

H. Graboske

L. Wong

September 15, 1977

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LAWRENCE LIVERMORE LABORATORY

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THE EQUATION OF STATE OF BERYLLIUM

Abstract

A new, wide-range equation of state (EOS) has been constructed for Be. The composite theoretical model incorporates ionization equilibrium and condensed-matter and multiphase physics. It also satisfies all thermodynamic equilibrium constraints. The theoretical EOS has been compared with all available high-pressure and high-temperature Be data, and satisfactory agreement is generally

achieved. The most interesting feature is the theoretical prediction of melting at just below 220 GPa (2 Mb), indicating an extremely wide pressure range for solid Be. A striking feature is the appearance of shell-structure effects in physical-process paths: 2 large loops appear on the principal Hugoniot and the behavior of release isentropes from $\rho = \rho_0$ is significantly affected.

Introduction

Beryllium is an important elemental metal used for a wide range of laboratory purposes, such as in LLL weapons programs. An hcp metallic solid with a high yield strength, it is unique as a low-Z metal. Its atomic structure and phase diagram are simple. The atomic structure is shown in Table 1. In the phase diagram, 1 the $\alpha(\text{hcp})$ phase undergoes a solid-solid transition to $\beta(\text{bcc})$ immediately prior

to melting at 1551 K.

The atomic structure of Be has some very interesting thermodynamic properties, as a result of two electron shells, equally populated but posessing a ten-fold difference in mean shell energy. The high first-ionization potential and the absence of a molecular state helps simplify the transition at lower densities from a multiphase metallic fluid to the plasma.

Table 1. The atomic structure of beryllium.

A	9.013	I _i (eV)	9.320	18.206	153.850	217.657
Z	4	Config.	1s ²	2s ²		
ρ _o	1.84 Mg/m ³					

Physics of the Primary EOS Codes

To calculate theoretical EOS, over the range of 1000-fold expanded density to 1000-fold compressed density from room temperature (0.02585 KeV) to 25 keV, we chose the best available theoretical model for each regime in the T. o range, and apply this model until it breaks down or becomes inaccurate. Typically, the entire equilibrium EOS surface must be handled as at least three separate regions: the ionization/dissociation equilibrium region (rigorous in the limit $T \rightarrow \infty$, or $\rho \rightarrow 0$), the condensed matter region (asymptotically exact as T \rightarrow 0 or $\rho \rightarrow \infty$), and the multiphase or chemical-physics region, which includes the near-normal density, low-temperature region. The theoretical models may be first-principal calculations (such as the coulomb gas-perturbation theory), asymptotic theories (Thomas-Fermi and other statistical atom models) or phenomenological models (normalized from experimental data).

The first difficulties inherent in this approach are encountered in constructing the composite EOS. The theoretical models may not interface well because they all tend to be asymptotically correct at widely differing extremes in the T, ρ range.

This means they all become inexact as they approach the "central" or overlapping regions of T, p range. Therefore, disagreements between models may tend to maximize just where they overlap. This leads to the deletion of portions of the EOS subspaces leaving gaps, called interpolation regions in the composite surfaces, where the various models must be joined numerically. The object of EOS theoretical studies is to reduce these gaps to a minimum and to eventually eliminate them completely.

More difficulties occur in the attempt to satisfy all fundamental thermodynamic constraints appropriate to equilibrium systems for the composite EOS. These constraints include the following conditions:

- All EOS functions must be continuous.
- Thermodynamic stability must be satisfied (C,, β_T positive).
- Thermodynamic consistency must
 be satisfied [(∂P/∂T)_V
 = (∂S/∂v)_T].

We monitor two additional constraints, called normality constraints since almost all "normal" equilibrium systems obey them. They are

$$(\partial P/\partial T)_{U} > 0$$
,

$$(\partial^2 P/\partial \rho^2)_T > 0$$
.

Meeting all of these restrictions, especially for the non-physical or numerical portions of the surface (the interpolation regions), is a demanding numerical task because of the complex temperature and density behavior of a modern quantum statistical mechanical EOS calculation. That is, the shell-structure regions (thermal and press) ionization zones, and dissociation zones) have complicated rapidly varying thermal structures as do the multiphase, chemical-physics regions with their large gradients and discontinuous derivative behavior.

For Be, the ionization-equilibrium (IE) region was computed with the OCCIPITAL code of Rouse, which calculates the statistical mechanics of partially ionized mixtures with a modified Saha method. Rouse has added first order coulomb corrections (the Debye-Hückel model) in a consistent manner to the Saha ionization equilibrium and the EOS model. In addition, he computes a complete excited-state sum using the Planck partition function. The model is the simplest extension of coulomb fluid theory

beyond the non-interacting, partial-ionization method of Saha. It is accurate over a somewhat wider T, ρ range than the ideal-plasma model, but becomes increasingly inaccurate with the onset of strong coulomb interactions, electron degeneracy, electron exchange, and correlation effects. The T, ρ range covered by calculations using this model is 0.5 to 2.5 × 10⁴ eV, 10^{-3} to 1.84 Mg/m³.

The condenged-matter portion of the Be EOS surface was computed using the TFCMIX code. 3 The finitetemperature Kirzhnits modification of Thomas-Fermi statistical atom theory, which includes exchange and quantum corrections to order h in the statistical atom framework, has been incorporated into this code. The ion fluid contribution to the thermodynamic properties is modeled by a plasma form, a Debye-Hükel weak screening limit, joined to a Brush-Sahlin-Teller Monte Carlo classicalion intermediate screening model, with a freezing transition at T = 21.6 to a harmonic Debye solid. This theoretical model was used in the T, p range from 0.02585 to 2.5 \times 10⁴ eV for $1.845 \le 0 \, (Mg/m^3) \le 2000$. It then interfaces with the IE EOS over the range 1.84 to 10 Mg/m^3 from 0.5 \leq T (eV) \leq 2.5 \times 10⁴.

The third region, multiphase, is the most difficult to compute in

Ā	9.012	T _m	1551 K
ρ	1.845 Mg/m^3	ΔS	9.637×10^{-5}
c	0.799 mm/µs (7.99 km/s)	8 _e	2.23×10^{-9}
s	1.124 cm/ ^µ s (11.24 km/s)	Ecoh	0.266 Mbar-cm ³ /g = $\left(2.4 \frac{\text{Mbar-cm}^3}{\text{mole}} \approx 26.6 \frac{\text{MJ}}{\text{Kg}}\right)$
	1.25		0.271
a	1.46	γe	2/3

that it requires a carefully selected set of input data and then a series of iterative calculations leading to a best match with all available experimental data. The theoretical vehicle for this procedure is the GRAY code, 4 a complex, multiphase equilibrium code designed for metals. The basic physics components of GRAY for this application include a Dugdale-McDonald form of the Gruneisen solid model, with the experimental Hugoniot data serving as the Gruneisen reference state. The electron contribution is represented by a modified free electron gas. To this is added a Lindemann-law melting transition and

a scaling-model hot-liquid regime. At lower densities, this is smoothly joined to a hard-sphere thuid model. The resultant system fleats the full three-phase region, with explicit treatment of the L-S, L-V, and S-V two-phase coexistence regions. The necessary input data used for Be are given in Table 2. The GRAY calculations were carried out over the range $10^{-3} \le \rho \, (\text{Mg/m}^3) \le 7 \text{ and } 0.02585$ $\leq T(eV) \leq 2$. This T, ρ subspace overlaps the TFC region in the range $0.02585 \le T(eV) \le 2 \text{ at } \rho = 7 \text{ Mg/m}^3$ and overlaps the low-T end of the IE region over the range $10^{-3} \le \rho \, (\text{Mg/m}^3)$ ≤ 1.845.

EOS Assembly

The assembly of these three subspaces into a single composite EOS is performed by various data-handling codes. The limits chosen for the various regions are rather arbitrary, a combination of estimates of model breakdown, plus the need to reach across to join to the other theoretical models. Since internal-validity criteria are not necessarily accurate in denoting theoretical inaccuracies, the degree and nature of disagreements between the various models at their interfaces can provide insight when

assessing the quality of the calculations. With this in mind, let us examine the interface details of the three surfaces.

An instructive representation for this purpose is the "thermal" energy surface. This is the internal-energy EOS surface plotted as a function of ρ with T as a parameter (Fig. 1). This thermal energy Eth is defined by

$$E_{th}(T_i, \rho_j) = E(T_i, \rho_j) - E(T_i, \rho_j)$$
,

where $E(T_1, \rho_i)$ is called the cold curve, meaning the lowest temperature isotherm on the EOS surface. subtraction of this component from the total surface remove the strong density dependence appearing at $\rho > c$ of the degenerate electron gas, where E $\propto \rho^{2/3}$. The resultant Eth surface has a "linearized" appearance for all T, ρ , since the remaining energy contributions are dominated by the classical ion and thermal electron components which are approximately linear in T. A comparison of Figs. 1 and 2 shows this graphically. The advantage of the energy surface over the pressure surface is that all of the thermal structure, such as ionization, pressure ionization, melt. and L-V coexistance regions, are more clearly identifiable.

The overlap between the condensed matter and multiphase regions is

shown in Fig. 1, where TFC data is used for $\rho > 10 \text{ Mg/m}^3$ and all T. while GRAY data extends from p ≤ 7 Mg/m³ for T ≤ 2 eV. The only modification required to the output EOS data was that the TFC energies be shifted down by 0.266 Mbar cm 3/gm (26.6 MJ/kg) because the TF calculations are relative to the free atom and, thus, must be corrected by the cohesive energy of the solid. When this is done, the TFC and GRAY pressure and energy surfaces join accurately, with the exception of a small region just above 2 eV near ρ_{o} where the TFC isotherms had to be deleted, appearing as a gap in Fig. 1. The excellent agreement between TFC and GRAY calculations is probably fortuitous. The fundamental differences in these two models usually require a renormalization of the entire TF region, a procedure less severe than it sounds, since the basic TF models are rigorously correct only at pressures of 1000 Mbar (100 TPa) and higher.

The second interface, between the multiphase and IE regions, can be seen in Fig. 1 also. Here the shell structure of the IE is evident, as is the multiphase structure, both melt and vaporization, of the chemical-physics region. The original juncture point shown here was the 2-eV isotherm, and it is clear from the figure this

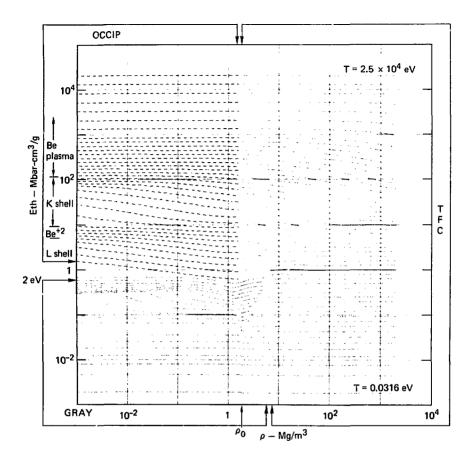


Fig. 1. The thermal energy surface, $E(T_1, \rho) - E(T_1, \rho)$, vs density for all T from 0.0316 to 2.5 × 10⁴ eV. The three theoretical subspaces are indicated.

was not adequate because the Be fourth electron was not fully recombined at 2 eV and low density. This association effect is evident from the decreasing energy of the lowest OCCIP isotherm (3 eV) when compared to the flat (ρ independent) ideal, neutral atomic-gas

behavior of the upper GRAY isotherm (2 eV). Because of this, a new interface, shown in Fig. 3., was chosen, and both models were compared over the range 0.5 to 2 eV. At that range, the OCCIP energies show that recombination of Be⁺ to a neutral Be atom

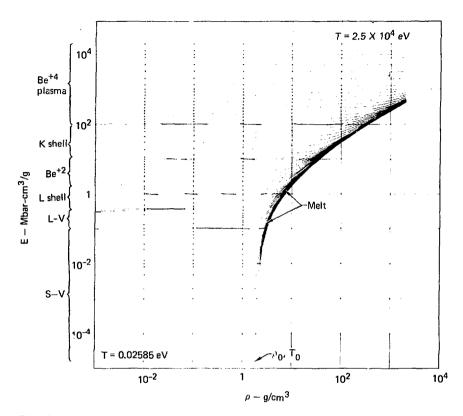


Fig. 2. The final Be energy EOS, plotted vs density. The basic T and ρ independent variable grid in 10 points per decade, approximately logarithmically spaced. This corresponds to the sequence 1, 1.25, 1.6, 2, 2.5, 3.16, 4, 5, 6.3, 8, for each decade, with $\rho_{\rm Ol}$ and 300 K (.02585 eV) as special values.

gas is only complete at 0.5 eV (for 10^{-3} to 1.845 Mg/m³). At this temperature GRAY is also predicting similar, ideal neutral-fluid Be energies (10^{-3} to 10^{-2} Mg/m³). The new OCCIP-GRAY interface is dropped to 0.5 eV, giving a smooth, accurate overlap of P and E at low density (10^{-3} to 10^{-2} Mg/m³). A problem does appear,

however: in examining this new interface for $\rho > 10^{-2}$ at low T, the non-ideal corrections to the ideal neutral Be gas become significant (at 0.5 and 0.63 eV), and, at about 0.1 Mg/m³, the GRAY model predicts the transition to the L-V two-phase condensation. Clearly, there is no way to join OCCIPITAL, which completely ignores

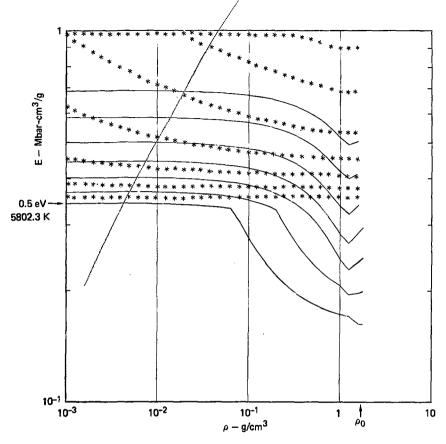


Fig. 3. The energy isotherms for T=0.5, 0.63, 0.8, 1, 1.25, 1.6 and 2 eV, as computed by OCCIPITAL (*) and GRAY (trace). The recombination of Be⁺ to Be is shown by the transition to a density independent OCCIPITAL isotherm as $T \rightarrow 0.5$ eV. The increase of non-ideal liquid corrections can be seen as departures from density-independent energy behavior in the GRAY isotherms, the shorp changes denoting onset of condensation.

these non-coulombic, liquid-like interactions, to the multiphase model in this region. The only resolution of this problem at present was to delete a portion of this surface and numerically reconstruct it (Fig. 4).

It corresponds to a regime where strong, liquid-like interactions are present with strong coulomb-plasma interactions, a physical system of extreme

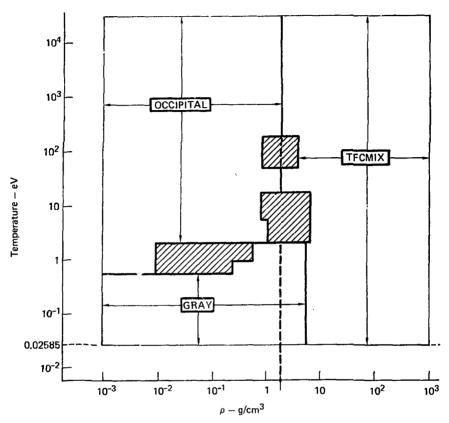


Fig. 4. The T, ρ independent variable space for the Be EOS calculations. The subspaces covered by the various theoretical codes are indicated, and the shaded regions denote the interpolation regions where adjustments have been made to interface the codes together.

complexity, which is currently theoretically intractable. The development of a dense plasma-chemical physics theory is clearly desirable to properly treat this region.

The third interface, between the IE and the TF regions, is fairly extensive, ranging from 2 eV to 25 keV,

at $\rho \approx \rho_0$. The shell structure in the 1E region of Fig. 1 shows a fully ionized Be plasma for T \geq 100 eV (at low density), a K-shell recombination zone for 100 eV to 30 eV. It also shows a Be⁺² "plateau" between roughly 30 eV and 12 eV, where the wide separation in K- and L-shell energies leaves

a stable Be⁺² plasma over a wide T. p range, and finally, it shows the recombination of the L shell, from 12 to 0.5 eV (the range 2 to 0.5 eV must be seen in Fig. 3 because it was removed in the Fig. 1 original subspace interface). Now when this shell structure is compared to the unstructured statistical atom (TFCMIX) isotherms along the ρ_0 isochore, several interesting features appear. First, for the ideal plasma (T ≥ 300 eV), the two theories agree exactly (as expected). But, as the coulomb interactions increase, they start to differ and the K-shell recombination produces a large discrepancy in thermal structure. A moderate agreement is again recovered in the Be⁺² plateau (probably fortuitous) until the onset of L-shell recombination and low-T deviations of the TF model destroy agreement. The net result of this is shown in Fig. 4, where a section of both OCCIP and TFC EOS surfaces has to be removed to accommodate the transition from structured K shell to pressureionized statistical atoms; a second larger segment is removed to accomo-The Be+2 date the L-shell transition. plateau region was directly joined, based on the apparent agreement of the two energy surfaces in the 12 to 30 eV region at ρ_0 . This choice was not a wise one, as shall be seen.

The final composite-EOS surface

in Fig. 4 is substantially represented by primary EOS code calculations, with perhaps 10% of the total T, p independent variable range replaced by numerical reconstruction. Along the various subspace interfaces the code to code agreement ranges from quite good to poor. It should be kept in mind here that these primary EOS codes may not be accurate, and a continual program to upgrade each theoretical area is necessary.

The numerical reconstruction of the interpolation regions is performed in a manner which attempts to satisfy the full set of thermodynamic constraints discussed above. A thermodynamically stable (TDS) method is used to fill the gaps, which ensures that the continuous EOS pressure and energy have positive derivatives C, and β_T . A second process is performed to renormalize the energy surface relative to the pressure surface so that the thermodynamic consistency (TDC) constraint is obeyed. If both these steps are successful, as was the case for Be, it appears that the thermodynamic normality (TDN) regulrements are also satisfied.

Two operational problems did occur in the TDS-TDC-TDN numerical processing. First, the L-V two-phase region had TDC problems, the result of numerical inaccuracies in the GRAY calculation. This will not have a major effect for most uses of the EOS,

but calculations which cross the L-V two-phase region should be critically examined. Second, when TDC integrations were performed to correct the energies in the interpolation regions, some TD instability (C_v negative) resulted in the 16- to 25-eV range for $\rho > 100 \text{ Mg/m}^3$. This was a consequence of the direct OCCIP-FFC interface made in the Be⁺² plateau region where the apparent good agreement in P and E of the two surfaces was not, in fact, adequate. This instability was removed, and the sensitivity of the TDC process to very small defects

in the pressure surface was recognized. It should be possible to use this sensitivity to locate regions where model interfaces are not sufficiently smooth and to iterate the interface-consistency operations to achieve a satisfactory result.

The final Be EOS energy is shown in Fig. 2. The various regions of special physical interest are indicated. This EOS and the pressure EOS are continuous, stable, consistent (with the exception noted above), and obey the TDN constraints.

EOS Analysis

The next step in the EOS-generation procedure is analysis of the EOS. The obvious step is a close comparison of the theoretical EOS with experimental data. A further useful step is to compare physical-process paths (Hugoniots, isentropes, isobars, etc.) over a large portion of the T, ρ range with other theoretical models.

There is a fairly small amount of high-quality experimental data on Be available for comparison with the theoretical EOS. The most useful data for EOS purposes is single-shock-generated, high-pressure Hugoniot data. We have chosen the Be data of McQueen et al. from Van Thiel's Compendium

of Shock Wave Data. 5 The single-shock data extends up to 885 kbar (88.5 GPa), and a comparison of theory and experiment is shown in Fig. 5. The good agreement between them is expected, in that the GRAY region uses a fit to the experimental $u_s - u_p$ data to define the compression $(\rho > \rho)$ portion of the multiphase region. It should be noted, though, that this theoretical Hugoniot is calculated from the final EOS primary data base (PDB) and is exactly the curve a hydrodynamic code would give using this EOS. A different form of the shock-wave data is shown in Fig. 6 (Hugoniot pressure versus density),

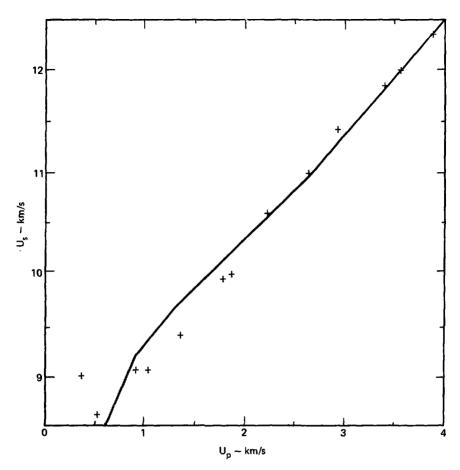


Fig. 5. Comparison of experimental and theoretical shock wave data, u_s vs u_p . The experimental data (*) of McQueen et al is very accurately reproduced by the theoretical Hugoniot (trace). The lowest point is probably an elastic precursor and should not be reproduced by the Hugoniot.

and in this space the agreement between theory and experiment is better. Beryllium shock data even at higher pressures would be desirable to check the theoretical model [885 kbar

(88.5 GPa) only corresponds to a density of 2.7 Mg/m^3].

A much more demanding test of the theoretical EOS can be made by a comparison with the single- and

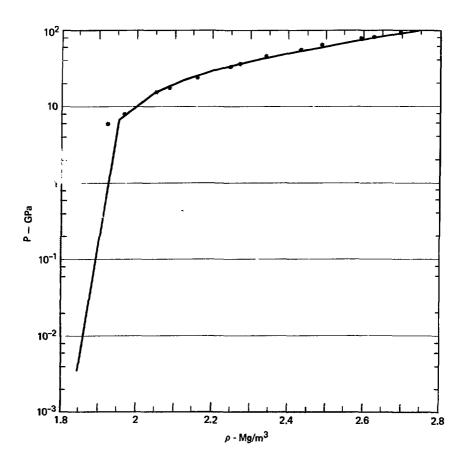


Fig. 6. Comparison of same experimental (*) and theoretical shock Hugoniot loci in pressure and density plane. Deviations in the 300 to 885 kbar range are of the order of a few per cent.

multiple-shock data of Neal.⁶ A series of experiments on Be were performed, using colliding shocks to reach the doubly and triply shocked states that lie between the principal

Hugoniot and the principal adiabat. The single-shock pressure achieved was 324 kbar (32.4 GPa), the double shock 788 kbar (78.8 GPa), and the convergent triply shocked state

1.41 Mbar (141 GPa). The latter two pressures were calculated from the model

$$\gamma_0 = 2S_0 - (t + 2)/3$$
,

using t = 0 (the Slater γ formulation). These results were compared to the Be theoretical EOS by computing single-. double-, and triple-shock states. can be seen from Fig. 7, the theoretical u_s-u_p curves for single, double and triple shocks pass quite close to the single- and doubleshock experimental data, and the error bars for the triple-shock data approach the curve. The same comparison in the pressure-density plane in Fig.. 8 shows how the double- and tripleshock points lie well below the principal Hugoniot. The substantially lower pressure of the triple-shock experiment relative to the theoretical triple Hugoniot indicates that the rate of decrease of Y with volume may be somewhat lower than the value chosen for the GRAY catculations.

The foregoing comparisons with experiments in the strongly compressed region of density only give information about the solid EOS model. If high-temperature data is examined, information concerning the melting and liquid behavior of the metal is elucidated. The best high-temperature data for Be is that discussed by Hultgren et al. They have selected an optimized set of heat-content data

extending up to 2200 K, with an extrapolation to 2800 K, primarily based on the work of Kantor et al. 8 Comparison between the enthalpy, $\Delta H \left(\frac{MJ}{k\sigma}\right)$ as a function of T at one atmosphere is shown in Fig. 9. The melting transition in the theoretical model is constrained to occur at the correct temperature; however, the melting density is modeled. Steinberg has noted a simple relation between the boiling temperature of elemental metals and their liquid density. From this relation, he estimates $\rho_n = 1.69$ (Mg/m³) for Be. The theoretical model predicts $\rho_n = 1.70$ and is in good agreement.

A further comparison in the high T, or expansion ($\rho < \rho_0$) part of the multiphase T, ρ range is given in Fig. 10. Here the melting curve for bcc Be is plotted from data in the compendium of Young. There is only moderate agreement between theory and experiment, since the theoretical melting temperatures are about 100 K low in the range 35-60 kbar (3.5-6.0 GPa). It should be kept in mind that, while errors of this magnitude are not desirable, the theoretical model is reasonably close to the true physical conditions.

A final overall comparison of the high-temperature plane can now be made by looking at three key high-T conditions: the melting, boiling and critical points. The first two conditions have been discussed; the critical

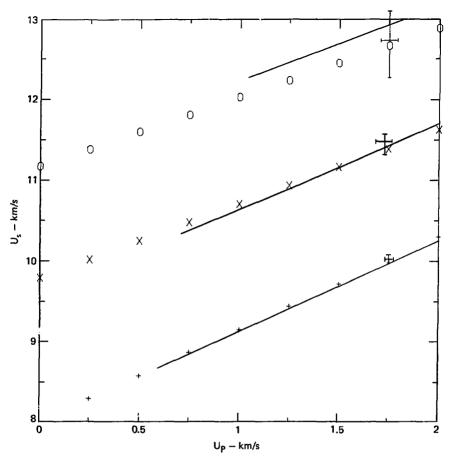


Fig. 7. Comparison of multiple shock experimental results of Neal with theoretical multiple Hugoniots calculated from the theoretical Be EOS. The experimental data points are shown (+) for single, double and triple shocks. Neal's $\mathbf{u_{S^{-\mathbf{u_{D}}}}}$ curves calculated from a Mie-Grüneisen EOS with a Slater γ are also shown (+, x, 0), while the theoretical multiple Hugoniot paths appear as traces. All are in reasonable agreement with the experimental points.

point for Be is unknown; however, using a method developed by Young and Alder 10 the critical conditions can be estimated. Using a Van der Waals model, a hard-sphere expression

from Monte Carlo computer experiments, and an effective hard-sphere estimate for the metal, Young estimates the critical conditions given in Table 3. It is clear from the table that the

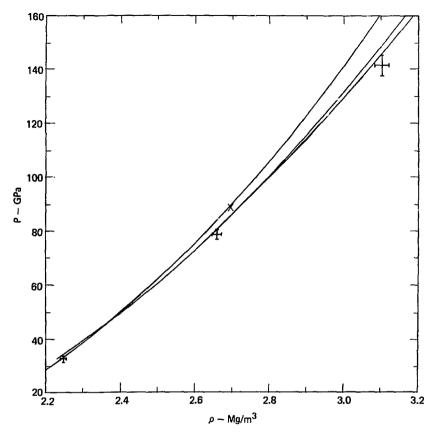


Fig. 8. Comparison of shock Hugoniot data with theoretical Hugoniots in the P-p plane. Neal's 32.4 GPa point, and the 88.5 GPa single shock point of McQueen et al hoth lie on the principal Hugoniot. Neal's double shock point lies just below (within experimental error) the reflected shock Hugoniot at 78.8 GPa. The triple shock point lies well below the triple shocked Hugoniot, but the large error bars make this point less critical for theoretical comparisons.

model matches the experimental data well near melt, but is less than satisfactory at the lower densities. This failure to match boiling and critical conditions is a consequence of the inadequacy of the simple hardsphere fluid model in representing the true Be dense-fluid regime. A close fit to either the boiling or critical conditions could have been

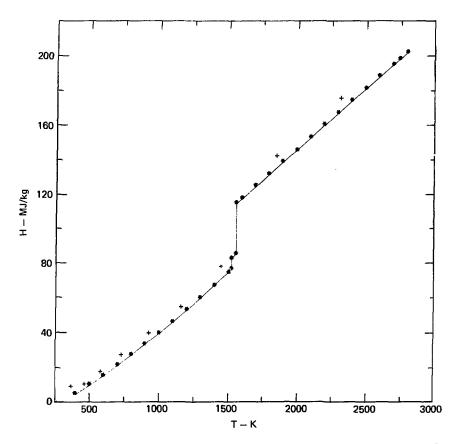
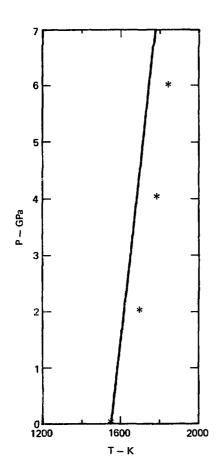


Fig. 9. The high temperature experimental data for heat content, $\Delta H (Mbar \cdot cm^3/g)$ versus temperature (K) at 1 atmosphere (101 KPa), taken from Hultgren et al (trace), compared to the theoretical 1 bar (101 KPa) isobar. The theoretical isobar exactly parallels the experimental one, with the melting transition occurring at 1551 K. The theoretical ΔH is equivalent to the solid-solid ($\alpha-\beta$) plus melting enthalpy change, since the theoretical model does not attempt to model the $\alpha-\beta$ transition.

achieved by appropriate adjustment of $E_{\rm coh}$. A choice of $E_{\rm coh}$ = 0.0377 MJ/mole (the true experimental value) would have exactly matched the boiling-point conditions, but with

a resulting T_c about two times larger than Young's estimate. Likewise, a much lower E_{coh} would give an exact match to critical-point estimates, at the cost of an even lower T_b than



that given in Table 3. The choice was made to adjust $\rm E_{coh}$ to come as close as possible to both $\rm T_b$ and $\rm T_c$. The correct alternative is to develop a more sophisticated dense-fluid model to allow accurate theoretical prediction of the entire liquid-vapor region.

All the foregoing discussion of experimental-theoretical agreement is limited to T \lesssim 0.25 eV (for actual data), a very low value compared to the full range of the EOS surface. The rest of the T, ρ range is devoid of any experimental checks. In the IE region, the Be ionization potentials are experimentally known, so the theoretical model is based on experiment as long as the ideal gas assumption remains valid, say, for $\rho < 10^{-2}$ Mg/m³, at T > 10 eV and $\rho < 10^{-1}$, T < 100 eV.

Fig. 10. The melting curve for Be. T_m versus P_m (bcc to liquid is shown for theoretical (trace and experimental (*) data. The experimental values are taken from Young's compendium, and extend to about 60 kbars (6 GPa).

Table 3. High temperature experimental data compared to the theoretical EOS.

Boiling			Melting		Critical Point		
Exper	imental	Theoretical	Experimental	Theoretical	Estimated	Theoretical	
T _B P _b	3090 K 1.366 Mg/m ³ 1 atm. (101) KPa)	2251 K 1.495 1 atm. (101 KPa)	T _m 1551 K ρ _m . 1.69 T _m —	1551 K 1.70 1971 K	T_c 7474 K ρ_c 0.493 P_c 1.22 × 10 ⁻² Mbar (1.22 GPa)	8484 K 0.606 1.71 × 10 ⁻²	
E _{coh}					(1.22 GFA)	(1.71 GPa)	

Similarly, at extremely high pressures, say P > 1000 Mbar (100 TPa), the TF models are strongly based on simple extensions of Fermi-Dirac electron statistics, also experimentally known. But it is evident that a wide gap from 1 Mbar to 1000 Mbar (0.1 to 100 TPa) or from 0.2 eV to 100 eV is totally based on the complicated theoretical models or on numerical thermodynamic approximations.

One final set of comparisons can be made to demonstrate the range of theoretical variance occurring in two different models. A well established theoretical model for Be is the work of Cowan, 11 who developed theoretical EOS for elements using finitetemperature, Thomas-Fermi-Dirac (TFD) physics for the electron gas, with ideal-gas ideal-solid ion contribu-These FOS were further cortions. rected by additional information from high-pressure shock data, limited to data obtained prior to 1957. Instead of the formidable task of comparing the full pressure and energy surfaces, it is more instructive to analyze certain key physical-process paths, such as Hugoniot paths and isentropic release paths. A comparison of the Cowan TFD-based Be theoretical Hugoniot with the corresponding track computed from the new Be EOS is shown in Fig. 11. The two theories agree well within the high-T

plasma limit [P > 5000 Mpar (500 TPa)]T > 1 keV] but differ significantly almost everywhere else. In the experimental regime $\{P \le 1 \text{ Mbar } (100 \text{ GPa})\}$, the older EOS does not have a defined Hugoniot: in the figure, the lowest Hugoniot pressure calculable is 1.55 Mbar (155 GPa), which results from its original design as a very highpressure equation of state. In the 1-1000 Mbar (0.1-100 TPa) intermediate region the major feature is the appearance of shell-structure (QM) ionization equilibrium, which causes the principal Hugoniots to differ even qualitatively. The TFD Hugoniot is relatively smooth and monotonic, averaged through the four ionization levels: the shell-structure Hugoniot exhibits two large loops, the lower one [4 to 30 Mbar (400 to 300 GPa)] occurring as the L-shell electrons are pressure ionized. There is a minimum between the shells, corresponding to the Be to plateau, then again a second loop, from 50 to 3000 Mbar (5 to 300 TPa), as the K-shell electrons are ionized. Be is a member of a small class of low-Z materials He, Li, Be and perhaps B, where the nearly equal population of the K and L shells gives comparable shell structure effects on the Hugoniot. The high-Z elements have pressure ionization effects smeared out over many electron shells, so their structure is much

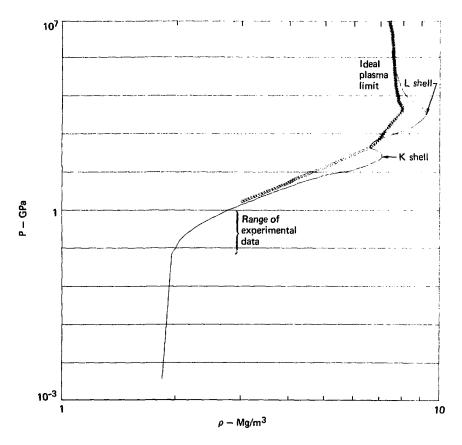


Fig. 11. The principal Hugoniots computed from two theoretical models. The TFD-based EOS of Cowan (***) shows typical smooth statistical atom behavior, while the new EOS goes from good agreement with shock data for P < 1 Mbar (100 GPa), through two shell structure pressure ionizations corresponding to the K and L shells.

closer to a smooth statistical atom model. Even for elements with Z of 6 to 12, the larger population of the L shell results in a smoothed L-shell structure, with only the K shell appearing in the pronounced manner seen here.

A second type of physical-path comparison is the release isentrope, where material is allowed to isentropically expand from some initial state. This is shown in Fig. 12 for the two theoretical EOS's. The initial states are 0.1, 1, 10, 100 and

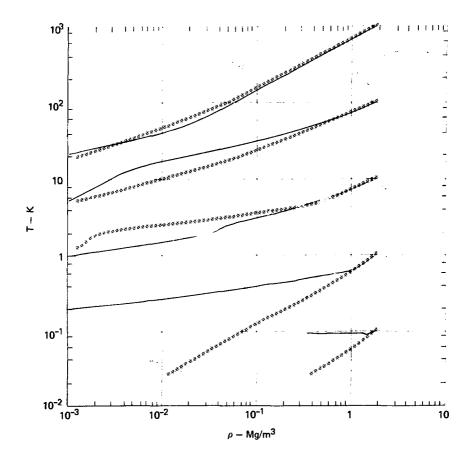


Fig. 12. Release isentropes in T, ρ space calculated from TFD (***) and the new Be EOS (trace). Significant differences exist in the expansion paths for all initial states from 0.1 to 1000 eV. The 10, 100, and 1000 eV paths pass through the low density thermal ionization regions, causing differences with statistical atom predictions, while the low T (0.1, 1 eV) paths pass through the dense neutral fluid and multiphase regions.

1000 eV, at $\rho=\rho_0$. In all the high-T cases, deviations between the TFD and shell-structure EOS are noted, including temperature differences as large as 80%. At these high T and

low ρ , the theoretical assumption of an ideal partial-ionization region is quite reasonable so these differences are probably real. This means equivalent differences will exist for P and E along these same tracks. Perhaps the major significance for these
differing T, ρ paths lies not in the
different hydrodynamic states attained, but in the use of the T values for calputating strongly T-dependent transport
properties, such as opacity.

using laser-generated shock waves
reach much higher pressures on the principal Hugoniot, 12 it might be interesting to see what shock
pressures would be required to me culating strongly T-dependent transport
Be and to reach the fully liquid state, according to the present

In the low-temperature regime (T ≤ 1 eV) the release isentropes differ tremendously, as would be expected. The TFD model of this regime, designed for very high pressure, yields a simple plasma-like release curve, new EOS gives very different behavior; the 1-eV isentrope stays in the fluid regime, passing just above the critical point. The 0.1 eV release path passes through melt and then into the two-phase region, where T is almost constant.

With the prospect of new techniques while the multiphase structure of the

using laser-generated shock waves to reach much higher pressures on the principal Hugoniot, 12 it might be interesting to see what shock pressures would be required to melt state, according to the present theoretical model. These values can be seen in Figs. 13 and 14 from the theoretical Hugoniot, and the melting and freezing curves. It appears that a single-shock pressure of about 1.95 Mbar (195 GPa) is needed to melt Be on the Hugoniot. This occurs at a temperature of 0.4 eV. To pass through the two-phase region and reach the fully liquid state, a 2.5 Mbar (250 GPa) shock is required, corresponding to 0.45 eV. It is quite possible that such shock strengths can readily be achieved by the existing large laser systems.

Summary and Conclusions

A complete theoretical EOS for Be has been generated, ranging from 10^{-3} to 2000 Mg/m³ and from 0.02585 to 2.5 × 10^5 eV. The basic theoretical models which have been combined to treat this material are a coulomb-corrected ionization equilibrium theory, OCCIPITAL; the finite-temperature, Thomas-Fermi-Kirzhnits electron model plus BST ion-fluid

contributions for the high-pressure, high-density regime, and the GRAY multiphase model, combining the Mie-Gruneisen solid, Lindemann law melting, scaling model liquid and hard sphere fluid. The resultant composite EOS is thermodynamically stable and consistent, and is very successful in matching the high-pressure shock Hugoniot data, both single and multiple

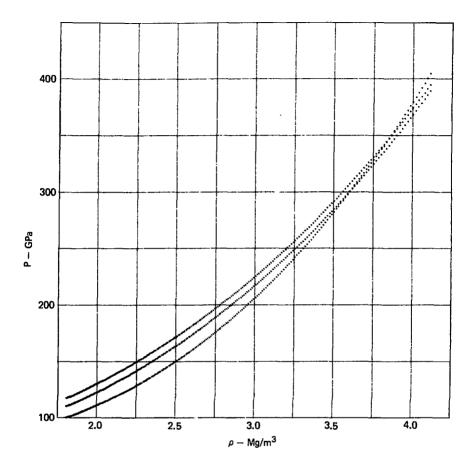


Fig. 13. The principal Hugoniot, melting (solidus) and freezing (liquidus) loci with pressure as a function of density calculated from the Be theoretical EOS. According to this work, the melting transition in Be could be observed at 1.95 Mbar (195 GPa), and complete transition to the liquid state would be achieved by 2.5 Mbar (250 GPa).

shock. It is only partially successful melting experimental data, and only in treating the high temperature data, roughly matches the boiling and esagreeing well with heat content and timated critical conditions. When

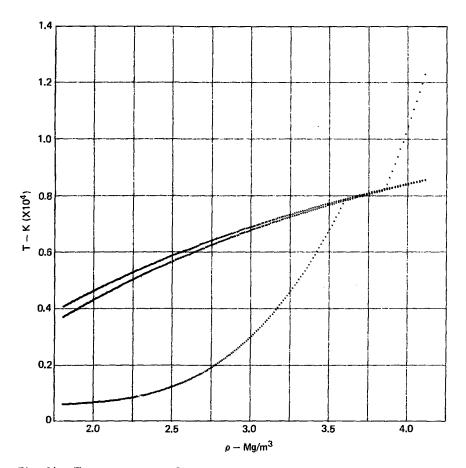


Fig. 14. The corresponding Hugoniot, melting and freezing loci in temperature density space, showing the onset of melting at 0.4 eV (η = 1.63), and complete liquefaction at 0.45 eV (η = 1.87).

compared to older pure statistical atom models, substantial differences are observed.

The basic features of the new Be

(1) Good agreement overall with

- a wide range of experimental data, both high pressure and high temperature.
- (2) The appearance of strong shell structure effects in the EOS surface, which have significant

sion and isentropic expansion from high T states.

effects on both shock compres- (3) A complete multiphase model of the S-L-V region, including melting and evaporation.

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